The KLOE Calorimeter Trigger

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KLOE aims to $\sigma(\Re e(\varepsilon'/\varepsilon)) \sim 10^{-4}$ with the double ratio method

$$\mathcal{R} = \frac{\Gamma(K_{\rm L} \to \pi^0 \pi^0)}{\Gamma(K_{\rm S} \to \pi^0 \pi^0)} \cdot \frac{\Gamma(K_{\rm S} \to \pi^+ \pi^-)}{\Gamma(K_{\rm L} \to \pi^+ \pi^-)} \simeq 1 - 6 \Re e \frac{\varepsilon'}{\varepsilon}$$

The detector has \sim 4 m diameter and \sim 4 m length:

- Superconducting Coil of 6 kG
- Lead-Scintillating Fiber Calorimeter with 13% energy sampling fraction and 5000 PMs readout



 \rightarrow the measured performances are :

$$\frac{\delta E}{E} = \frac{5.7\%}{\sqrt{E(GeV)}}; \quad \delta t = \frac{54ps}{\sqrt{E(GeV)}} + 147ps$$

• Helium Drift Chamber all-stereo cell geometry with 12500 sense wires

• Equalized trigger efficiencies on $\pi^+\pi^-$ and $\pi^0\pi^0$ decays of $K_{\rm L}$ and $K_{\rm S}$:

$$\rightarrow \mathcal{R} = \frac{N_L^0 N_S^{\pm}}{N_S^0 N_L^{\pm}} \cdot \left(1 + \varepsilon_S^0 - \varepsilon_S^{\pm} + \varepsilon_L^{\pm} - \varepsilon_L^0\right)$$

• High efficiency \rightarrow small $\sigma(\varepsilon) = \sqrt{\varepsilon(1-\varepsilon)/N}$



Phi physics rate: ~2 kHz



Bhabha rate: 3.5 kHz



Cosmic rays rate: 2.6 kHz



Machine Background rate: O(MHz)

- DAQ max throughput: 50 Mbytes/sec (\sim 10 kHz) \rightarrow rejection capability on "non- ϕ events"
- 2.7 ns interbunch spacing \rightarrow continuous mode operation

• Two independent chains:

1) EmC trigger: based on multiplicity of fired calorimeter sectors

2) DC trigger: based on multiplicity of hit drift chamber wires

• Two level scheme:

1) Fast L1 (within 200 ns) to start calorimeter TDC's

2) Slower L2 confirmation after $\simeq 2\mu$ s (drift time in the chamber cells)



• L1 synchronized with machine RF/4 and distributed with 50 ps precision

Two energy deposits are required in a 70 ns time window.

Each sector discriminates signals on two thresholds at the same time:

• A low threshold (LET = 50 MeV on barrel, 85 MeV on the endcaps), in order to trigger on particles from ϕ decays in the following topologies



• A high threshold (BBT = 300 MeV) in order to identify and reject/downscale Bhabha events



Bhabha event on barrel



Each trigger channel discriminates the two input signals, coming from side A and B of the calorimeter with 8 comparator:
4 produce the LET signal, and 4 the BBT one.



DISH scheme



The threshold profile is measured by correlating the reconstructed sector energy with the corresponding trigger response, in z slices. The threshold spread along z is compatible with the energy resolution in the 40-100 MeV range (30% - 20%)





A calibration curve has been obtained spanning the range of all possible DISH settings

Dish low LET (counts)



 $\rightarrow\,$ DISH boards provide signals proportional to the number of active trigger channels

 \rightarrow PASTA board uses these signals to count the number of " ϕ tagged" (LET) and "Bhabha tagged" (BBT) particles, separately for Barrel and Endcaps, merging the information coming from two series of overlapping trigger sectors.

1) PIZZA sum stage: calorimeter ADC signals versus the trigger ADC at DISH inputs 2) Trigger sector's hardware: LET occupancy (blue) versus the illumination of calorimeter cells (red)



3) Consistency check on the multiplicity counting stage: multiplicities at the TORTA input are correlated with the ones calculated starting from DISH outputs



BHABHA SELECTION:

- 2 BBT sectors on the barrel;
- Angular distance: $15 \le \Delta \phi \le 22$ trigger sectors;
- Time of flight difference on trigger TDC's: $\Delta t \leq 5$ ns.



PERFORMANCES:

- Fast and reliable response \rightarrow one luminosity measurement per minute at $10^{31}~{\rm cm}^{-2}{\rm s}^{-1}$ with $\sigma_{STAT}=5\%$;
- $\mathcal{L}_{TRIG} \mathcal{L}_{OFFLINE} \sim 10\%$;
- Rejection power on cosmics $> 10^5$ (~ 4 mHz residual rate);
- Machine background contamination $\leq 5\%$.

The distribution of clusters with E $>100~{\rm MeV}$ shows that machine background peaks around the beam pipe, particularly toward the centre of DA $\Phi{\rm NE}$ orbit



Bhabha trigger channels on the endcaps (veto not needed if ${\cal L} < 10^{32}~{\rm cm^{-2}s^{-1}}$) are currently used to monitor single rates with a 100 MeV threshold





Close to the beam pipe higher thresholds have been set in order to reject accidental coincidences of the two endcaps

In the present running condition ($I^-/I^+ \simeq 500$ mA; $\mathcal{L} \simeq 10^{31}$ cm⁻²s⁻¹) the unbalanced threshold map reduces the trigger rate from ~ 3000 Hz down to ~ 1300 Hz (calorimeter trigger only), composed by:

a) **650 Hz** of cosmic rays, b) **70 Hz** of Bhabha's, c) **30 Hz** of ϕ decays + d) **550 Hz** machine background



 ${\sf K}_{\rm S} \to \pi\pi$ events are tagged by a ${\sf K}_{\rm L}$ interacting in the calorimeter barrel with E > 100 MeV.

 $K_{\rm L}$ fires at least 1 sector, the trigger efficiency is therefore (2 sectors required !):

$$\varepsilon(\mathbf{K}_{\mathrm{S}} \to \pi\pi, \mathbf{K}_{\mathrm{L}} \mathrm{crash}) = 1 - \mathbf{P}_{\mathrm{L}}(1) \cdot \mathbf{P}_{\mathrm{S}}(0)$$



The following efficiencies have been measured for $K_{\rm L}$ and $K_{\rm S}$:

$$\begin{aligned} \mathbf{P}_{\rm S}(0) &= 0.24 \pm 0.03\% \ (\pi^0 \pi^0) \\ \mathbf{P}_{\rm S}(0) &= 2.69 \pm 0.07 \ \% \ (\pi^+ \pi^-) \end{aligned} \qquad \mathbf{P}_{\rm L}(1) = 55.0 \pm 0.3 \ \% \end{aligned}$$

As a result:

- $\epsilon_{trig} = (99.87 \pm 0.02) \%$ for $\pi^0 \pi^0$ $\epsilon_{trig} = (98.52 \pm 0.04) \%$ for $\pi^+ \pi^-$

 The probability P(i) of having "i" sectors fired (i = 0, 1, 2) on each subdetector (B = barrel, W = endcap west, E = endcap east) is:

$$P(0) = S(0)L(0); P(1) = S(0)L(1) + S(1)L(0), \dots$$

where S(i) and L(i) are extracted from the "unbiased" multiplicity distributions of ${\rm K}_{\rm S}$ and ${\rm K}_{\rm L}$

• The fraction of lost events 1 - ϵ_{trig} is:

$$B(0)W(0)E(0) + B(0)W(1)E(0) + B(0)W(0)E(1)$$

For
$$K_{\rm L}
ightarrow 3\pi^0$$
 (tagged by $K_{\rm S}
ightarrow \pi^+\pi^-$)

$$\epsilon_{trig} = 99.0 \pm 0.5\%$$



- The KLOE calorimeter trigger selects with high efficiency the neutral kaon decays
- It also produces a fast signal for starting the calorimeter FEE within 300 ns from the interaction
- Despite high backgrounds, the S/B ratio is kept at a reasonable level
- A highly redundant monitoring system allows for an efficient online control of the detector hardware, providing at the same time reliable measurements of the machine luminosity and the background levels

$BHABHA_VETO = B2B + E1B * W1B$

 \rightarrow Efficiency $\sim 96\%$.



 $COSMIC_VETO = B2C + B1C * (E1C + W1C)$

 \rightarrow Efficiency $\sim 75\%$ (2.7 kHz \rightarrow 0.68 kHz)



The ~25% cosmic veto inefficiency is dominated by the geometry of the apparatus



Trigger Scheme



Trigger timing



Rephasing

